Home Search Collections Journals About Contact us My IOPscience

Strong-coupling expansions for the  $\mathcal{PT}$ -symmetric oscillators  $V(\mathbf{\textit{x}}) = \mathbf{\textit{a}}(\mathbf{i}\mathbf{\textit{x}}) + \mathbf{\textit{b}}(\mathbf{i}\mathbf{\textit{x}})^2 + \mathbf{\textit{c}}(\mathbf{i}\mathbf{\textit{x}})^3$ 

This article has been downloaded from IOPscience. Please scroll down to see the full text article. 1998 J. Phys. A: Math. Gen. 31 10105 (http://iopscience.iop.org/0305-4470/31/50/008)

View the table of contents for this issue, or go to the journal homepage for more

Download details: IP Address: 171.66.16.104 The article was downloaded on 02/06/2010 at 07:22

Please note that terms and conditions apply.

# Strong-coupling expansions for the $\mathcal{PT}$ -symmetric oscillators $V(x) = a(\mathbf{i}x) + b(\mathbf{i}x)^2 + c(\mathbf{i}x)^3$

F M Fernández<sup>†</sup>, R Guardiola<sup>‡</sup>, J Ros<sup>§</sup> and M Znojil

† CEQUINOR, (Conicet, UNLP), Facultad de Ciencias Exactas, Universidad Nacional de La Plata, Calle 47 entre 1 y 115, Casilla de Correo 962, 1900 La Plata, Argentina
‡ Departamento de Física Atómica y Nuclear, Universidad de Valencia, Avda. Dr. Moliner 50, 46100-Burjassot, Valencia, Spain
§ Departamento de Física Teórica and IFIC, Universidad de Valencia, Avda. Dr. Moliner 50, 46100-Burjassot, Valencia, Spain

|| Oddělení teoretické fyziky, Ústav jaderné fyziky AV ČR, 250 68 Řež u Prahy, Czech Republic

Received 7 October 1998

**Abstract.** We study the traditional problem of convergence of perturbation expansions when the hermiticity of the Hamiltonian is relaxed to a weaker  $\mathcal{PT}$  symmetry. An elementary and quite exceptional cubic anharmonic oscillator is chosen as an illustrative example of such models. We describe its perturbative features paying particular attention to the strong-coupling regime. Efficient numerical perturbation theory proves suitable for such a purpose.

## 1. Introduction and summary

The analysis of convergence of perturbation expansions is one of the most exciting mathematical problems in quantum physics. Its relevance ranges from the phenomenology of bound states in atomic and molecular physics up to abstract methodical considerations which concern quantum fields [1]. Often, the convergence of perturbation series is studied by means of an elementary anharmonic oscillator model in one dimension:

$$V(x) = B x^{2} + C x^{3} + D x^{4} \qquad x \in (-\infty, \infty)$$
(1)

because of its apparent simplicity. The application of textbook Rayleigh–Schrödinger (RS) perturbation theory is remarkably simple in the so called weak coupling regime dominated by harmonic oscillations when  $B \gg |C|$  and  $B \gg |D|$ .

In more sophisticated approaches based on an extensive numerical re-processing of the RS series one considers the strong-coupling regime in which the quartic term is not only important but also dominating [2, 3]. This is of particular importance in field theory where perturbation expansions are one of the few available constructive means. The strong-coupling approach may also be of value in a quantitative description of seemingly non-perturbative quantum mechanical systems of physical interest [4]. Allowing complex values of the couplings of model (1) further enriches its physical meaning and scope [5]. It opens new phenomenological possibilities. Energies may acquire purely imaginary components, bound states may 'dissolve' into unstable resonances, etc. Various mathematical techniques have to be combined in order to treat these extended models with non-Hermitian Hamiltonians (see [6] for a review of some relevant literature).

0305-4470/98/5010105+08\$19.50 © 1998 IOP Publishing Ltd

10105

## 10106 F M Fernández et al

In particular, Bender and Boettcher [7] recently discovered, among all the complex quartic potentials (1), a large two-parametric family of non-Hermitian forces which are partially, *quasi-exactly* [8] solvable. The parity-violating Bender and Boettcher example is already slightly counterintuitive and exceptional. Its appeal comes from field theory as it conserves the product  $\mathcal{PT}$  of parity and time reversal. In the context of quantum mechanics this just means the symmetry with respect to the simultaneous reflection of the coordinate  $x \rightarrow -x$  and complex conjugation which, formally, replaces i by -i [9].

Strictly speaking, a rigorous treatment of quartic potentials and, in general, of all their higher-power  $\mathcal{PT}$ -symmetric generalizations  $(ix)^N$  with  $N \ge 4$  already requires a nontrivial deformation of the integration path into the complex plane [10]. A slight formal simplification is achieved, therefore, when the dominant quartic anharmonicity in (1) vanishes [11]. In this setting, trying to bridge the gap between the Hermitian and non-Hermitian Hamiltonians in what follows, we will analyse only the non-Hermitian and  $\mathcal{PT}$ -symmetric shifted version of (1) with D = 0:

$$V(x) = a(ix) + b(ix)^{2} + c(ix)^{3} \qquad x \in (-\infty, \infty).$$
(2)

Historically, the empirical and numerically oriented studies of interactions (2) with a = 0 go back to Daniel Bessis who conjectured, several years ago, that the purely imaginary cubic coupling keeps the 'resonant' anharmonic energies discrete and, what is even more surprising, *purely real* [12]. This property makes these models immediately eligible for perturbative description.

A new, additional motive for perturbative analysis of non-Hermitian oscillators of Bessis type lies in unexpected difficulties related to their weak coupling perturbative interpretation. Indeed, one cannot allow the cubic *and* quadratic coupling to vanish simultaneously since the spectrum of linear  $V_0(x) = ix$  is null [13]. In this sense, besides a far reaching analogy between equations (1) and (2) (which was the original inspiration for our present paper) there also exist certain differences.

Several particular cases of potential (2) with real couplings are discussed here. In section 2 we recall the ideas of numerical perturbation theory [14, 15]. Section 3 then shows that their application to the class of non-Hermitian examples (2) is straightforward. Finally, in section 4, we add more observations regarding re-summations of perturbative expansions for our  $\mathcal{PT}$  symmetric Hamiltonians, motivated by their possible methodical connection to field theory, etc.

We may summarize by stating that even our utterly schematic examples confirm that many  $\mathcal{PT}$ -symmetric non-Hermitian oscillators similar to (2) may be understood and described in a way which strictly parallels the existing extensive studies of the ordinary anharmonic oscillator (1).

The continuation of real Hamiltonians to the complex plane while preserving their  $\mathcal{PT}$  symmetry opens a new and almost unexplored field of mathematical analysis of Schrödinger equation. In this context we have resorted to the simple numerical algorithm to compute perturbation series. For Hermitian Hamiltonians, such a numerical form of perturbation theory proved convenient as a stable source of expansions that are suitable for all values of the coupling constant. Our present results extend this numerical experience to a few extremely interesting non-Hermitian examples.

The numerical perturbative approach is again shown to offer a reliable computational tool. In the strong-coupling and, possibly, renormalized regime the possibility of using a virtually arbitrary zero-order system of Bessis type is well matched by the 'perturbation friendly' real and discrete character of the spectra of its  $\mathcal{PT}$ -symmetric perturbations. Of course, the eigenfunctions are complex valued, with the only requirement being that of

giving square-integrable wavefunctions. One may appreciate the role of the  $\mathcal{PT}$  symmetry of Hamiltonians which not only provide *real* eigenvalues, but also their perturbation form with *real* coefficients.

#### 2. Perturbation series by numerical techniques

Traditional textbooks on quantum mechanics pay a thorough attention to the construction of perturbative expansions based on the expansion in a complete set of eigenvectors of the unperturbed Hamiltonian  $H_0$ . For this reason, the perturbation series

$$E = E_0 + \lambda E_1 + \lambda^2 E_2 + \cdots$$
(3)

for the eigenvalues of Hamiltonians  $H(\lambda) = H_0 + \lambda W$  are commonly restricted to a small range of a few available exactly solvable models.

In contrast, a purely numerical approach is both straightforward and more widely applicable to the determination of perturbation expansions like (3). A simple and efficient procedure for obtaining RS perturbation expansions for the solutions of the Schrödinger equation with Hamiltonians

$$H = -\frac{\mathrm{d}^2}{\mathrm{d}x^2} + V(x) + \lambda W(x) \tag{4}$$

consists in a replacement of the second derivative operator by the centred second difference operator,

$$\frac{\mathrm{d}^2}{\mathrm{d}x^2} \approx \frac{\delta^2}{h^2}$$

h being the distance between adjacent points in a suitable uniform mesh. In principle, the problem reduces to the diagonalization of a *symmetric* and *tridiagonal* matrix, well known and understood in the case of real potentials [16].

Many open questions regarding numerical methods for complex potentials still exist [17]. In particular, our Hamiltonian matrices cease to be Hermitian even after discretization. One must test and verify the very reality of their eigenvalues, as well as the rate of their  $h \rightarrow 0$  convergence with much greater care. A word of warning may come from both the elementary analytical arguments (illustrating, for example, the subtleties of discretization in the complex plane by the—very discontinuous—square well example) and the numerical tests we offer here.

From this point on, upper-case bold sans-serif and lower-case bold italic letters denote square matrices and column vectors, respectively. In such a notation we solve the matrix eigenvalue problem

$$[\mathbf{H}_0 + \lambda \mathbf{H}_I] \boldsymbol{v} = \boldsymbol{E} \boldsymbol{v}$$

where the Hamiltonian matrix  $\mathbf{H}_0 + \lambda \mathbf{H}_I$  is symmetric, with

$$\mathbf{H}_{0} = \begin{bmatrix} \frac{2}{h^{2}} + V_{0} & -\frac{1}{h^{2}} & 0 & 0 & \cdots \\ -\frac{1}{h^{2}} & \frac{2}{h^{2}} + V_{1} & -\frac{1}{h^{2}} & 0 & \cdots \\ 0 & -\frac{1}{h^{2}} & \frac{2}{h^{2}} + V_{2} & -\frac{1}{h^{2}} \\ \vdots & \ddots & \ddots & \ddots \end{bmatrix}$$

10108 F M Fernández et al

and

$$\mathbf{H}_{I} = \begin{bmatrix} W_{0} & 0 & \cdots \\ 0 & W_{1} & \\ \vdots & & \ddots \end{bmatrix}.$$

When applying perturbation theory one first solves the eigenvalue equation for the unperturbed problem

$$\mathbf{H}_0 \boldsymbol{v}_0 = \boldsymbol{E}_0 \boldsymbol{v}_0 \tag{5}$$

and then a hierarchy of inhomogeneous equations for the perturbation corrections of higher order:

$$\mathbf{H}_{0}v_{k} + \mathbf{H}_{I}v_{k-1} = E_{0}v_{k} + E_{k}v_{0} + \sum_{p=1}^{k-1} E_{p}v_{p-k}.$$
(6)

Notice that  $v_0$  is both a left and right eigenvector because of the symmetry of  $\mathbf{H}_0$ . We find it convenient to use the scalar product  $[v^Tw] \equiv \sum v_i w_i$ , although it does not produce a real norm, and resort to the *intermediate normalization* condition

$$[\boldsymbol{v}_{\mathrm{o}}^{\mathrm{T}}\boldsymbol{v}_{k}]\equiv\sum_{i}v_{\mathrm{o}i}v_{ki}=\delta_{k\mathrm{o}i}$$

that leads to simpler expressions for the energy coefficients. The intermediate normalization condition is based on the observation that if  $v_k$  is a solution of the perturbation equation, then  $v_k + \alpha v_0$  is also a solution for arbitrary values of  $\alpha$ . The computational scheme is straightforward:

(i) Solve the eigenvalue equation (5).

(ii) Project (6) onto  $v_0^{\rm T}$  and obtain the kth perturbative correction to the energy

$$E_k = [v_0^{\mathrm{T}} \mathbf{H}_I v_{k-1}].$$

(iii) Once  $E_k$  is known, solve the inhomogeneous equation for  $v_k$ .

(iv) Orthogonalize  $v_k$  with respect to  $v_0$ :

$$v_k \leftarrow v_k - ig[v_{ ext{o}}^{ ext{T}} v_kig] v_{ ext{o}}$$

## 3. Application to the two $\mathcal{PT}$ symmetric models

Table 1 shows some perturbation corrections to the energies of the  $\mathcal{PT}$  symmetric Hamiltonians

$$H^{(1)} = p^2 + ix^3 + i\lambda x$$
(7)

and

$$H^{(2)} = p^2 + ix^3 + \lambda x^2.$$
(8)

In both cases the coefficients  $E_n^{(1,2)}$  of the powers of  $\lambda$  are *real*, as expected from the  $\mathcal{PT}$  symmetry. We used complex arithmetics for the numerical calculation observing that the imaginary parts vanished within the rounding errors. The code was successfully tested with the exactly solvable model

$$H = p^2 + x^2 + i\lambda x \tag{9}$$

obtaining the known expansion

$$E(\lambda) = 1 + \frac{1}{4}\lambda^2$$

which one can easily derive by means of the coordinate translation  $x \to x + i\lambda/2$  in the Hamiltonian  $H = p^2 + x^2$ . The agreement between the numerical and exact result suggests that the method is sufficiently accurate and stable.

Table 1. Energy coefficients for the ground states of the Hamiltonians (7) and (8).

| n  | $E_n^{(1)}$ for $p^2 + ix^3 + i\lambda x$ | $E_n^{(2)}$ for $p^2 + ix^3 + \lambda x^2$ |
|----|---|--|
| 0  | 1.156 267 071 982                         | 1.156 267 071 982                          |
| 1  | 0.590 072 533 078                         | 0  |
| 2  | 0.119 413 858 091                         | 0.196 690 844 356                          |
| 3  | $-1.142128489951\times10^{-2}$            | $-7.407407406938 	imes 10^{-2}$            |
| 4  | $2.197\ 147\ 795\ 152	imes 10^{-3}$       | $1.326820645262 	imes 10^{-2}$             |
| 5  | $-5.299287434734 	imes 10^{-4}$           | 0  |
| 6  | $1.432~651~458~960 	imes 10^{-4}$         | $-4.230105523351 	imes 10^{-4}$            |
| 7  | $-4.148903908710	imes10^{-5}$             | 0  |
| 8  | $1.257939817338 	imes 10^{-5}$            | $2.712528133232 \times 10^{-5}$            |
| 9  | $-3.941\ 134\ 497\ 194	imes 10^{-6}$      | 0  |
| 10 | $1.265478160326 	imes 10^{-6}$            | $-2.180776735057 \times 10^{-6}$           |
| 11 | $-4.141789581668 \times 10^{-7}$          | 0  |
| 12 | $1.376462051952 	imes 10^{-7}$            | $1.965228322941 	imes 10^{-7}$             |
| 13 | $-4.632126852199 	imes 10^{-8}$           | 0  |
| 14 | $1.575\ 200\ 873\ 072	imes10^{-8}$        | $-1.897075421079	imes10^{-8}$              |
| 15 | $-5.404\ 290\ 239\ 370	imes10^{-9}$       | 0  |
| 16 | $1.868293677263	imes 10^{-9}$             | $1.917\ 298\ 897\ 845	imes 10^{-9}$        |
| 17 | $-6.501\ 605\ 918\ 143 \times 10^{-10}$   | 0  |
| 18 | $2.275\ 683\ 198\ 987	imes10^{-10}$       | $-2.002\ 303\ 776\ 843 	imes 10^{-10}$     |
| 19 | $-8.006218039652 	imes 10^{-11}$          | 0  |
| 20 | $2.829596711623 \times 10^{-11}$          | $2.143\ 098\ 365\ 950 \times 10^{-11}$     |

The two perturbative expansions of table 1 are connected, because the Hamiltonians (7) and (8) are related by a complex translation of the coordinate. The substitution  $x \rightarrow x + i\lambda/3$  in (8) results in a Hamiltonian like equation (7):

$$H \rightarrow p^2 + ix^3 + \left(\frac{i\lambda^2}{3}\right)x - \frac{2}{27}\lambda^3$$

Consequently, the energy coefficients  $E_n^{(1)}$  and  $E_n^{(2)}$  of table 1 satisfy

$$E_{2n}^{(2)} = \frac{1}{3^n} E_n^{(1)} \qquad E_{2n+1}^{(2)} = -\frac{2}{27} \delta_{n1}$$
(10)

indicating that the expansion  $E_n^{(2)}$  has only even coefficients except for  $E_3^{(2)} = -\frac{2}{27}$ . Our numerical coefficients shown in table 1 obey the exact relations (10) with an accuracy close to 1 part in  $10^{10}$ . This test is an additional confirmation of the stability and accuracy of our numerical method.

We may conclude that the application of a numerical version of perturbation theory to non-Hermitian Hamiltonians is straightforward as long as their spectrum is discrete. The demonstrated feasibility of a reliable *quantitative* perturbative description of our unusual models (2) with  $\mathcal{PT}$ -symmetric forces is encouraging.

## 4. Renormalized perturbation expansions

From the magnitude of the energy coefficients in the tests of section 3 we may estimate the radii of convergence to be about 1.7 and 3 for the Hamiltonians (7) and (8), respectively.

|    | variable (1                     | $(x), 101 \ \omega =$ | i, while the factor [x/(i       | <i>x</i> )] . |                                  |
|----|---------------------------------|-----------------------|---------------------------------|---------------|----------------------------------|
| n  | $E_n^{(r)}$                     | n                     | $E_n^{(r)}$                     | n             | $E_n^{(r)}$                      |
| 0  | $1.15626707 	imes 10^{+0}$      | 30                    | $-6.17135480 	imes 10^{-4}$     | 60            | $-5.42186148 	imes 10^{-5}$      |
| 1  | $-4.62506829	imes10^{-1}$       | 31                    | $-8.97409509\times10^{-4}$      | 61            | $-1.39700206 	imes 10^{-5}$      |
| 2  | $5.79387957	imes 10^{-2}$       | 32                    | $-1.05471829	imes10^{-3}$       | 62            | $2.19313729 	imes 10^{-5}$       |
| 3  | $8.79538466	imes 10^{-2}$       | 33                    | $-1.09662861	imes10^{-3}$       | 63            | $5.12970044 	imes 10^{-5}$       |
| 4  | $7.66512627	imes10^{-2}$        | 34                    | $-1.03884674 	imes 10^{-3}$     | 64            | $7.274\ 120\ 89 	imes 10^{-5}$   |
| 5  | $5.723~695~34 	imes 10^{-2}$    | 35                    | $-9.02533600 	imes 10^{-4}$     | 65            | $8.56777888 	imes 10^{-5}$       |
| 6  | $3.81037038\times10^{-2}$       | 36                    | $-7.11680394	imes10^{-4}$       | 66            | $9.025\ 593\ 62 	imes 10^{-5}$   |
| 7  | $2.18853370	imes10^{-2}$        | 37                    | $-4.90726498 	imes 10^{-4}$     | 67            | $8.724\ 869\ 94 	imes 10^{-5}$   |
| 8  | $9.294~671~98 	imes 10^{-3}$    | 38                    | $-2.62552992 	imes 10^{-4}$     | 68            | $7.79089589 	imes 10^{-5}$       |
| 9  | $2.69486228	imes10^{-4}$        | 39                    | $-4.69358453 	imes 10^{-5}$     | 69            | $6.380\ 829\ 04	imes 10^{-5}$    |
| 10 | $-5.59041145	imes10^{-3}$       | 40                    | $1.40504096	imes 10^{-4}$       | 70            | $4.667~329~11 \times 10^{-5}$    |
| 11 | $-8.82442831 	imes 10^{-3}$     | 41                    | $2.88858970	imes 10^{-4}$       | 71            | $2.82321647 	imes 10^{-5}$       |
| 12 | $-1.00084522 	imes 10^{-2}$     | 42                    | $3.92038967 	imes 10^{-4}$      | 72            | $1.008\ 185\ 15	imes 10^{-5}$    |
| 13 | $-9.69724568 	imes 10^{-3}$     | 43                    | $4.48490339 	imes 10^{-4}$      | 73            | $-6.41694726 	imes 10^{-6}$      |
| 14 | $-8.39124475 	imes 10^{-3}$     | 44                    | $4.60609691	imes10^{-4}$        | 74            | $-2.02126036 \times 10^{-5}$     |
| 15 | $-6.51827989	imes10^{-3}$       | 45                    | $4.33946792 	imes 10^{-4}$      | 75            | $-3.060\ 201\ 07 \times 10^{-5}$ |
| 16 | $-4.42531487	imes10^{-3}$       | 46                    | $3.76289089 	imes 10^{-4}$      | 76            | $-3.723\ 809\ 71 \times 10^{-5}$ |
| 17 | $-2.37729832 	imes 10^{-3}$     | 47                    | $2.96715278	imes10^{-4}$        | 77            | $-4.01090885 	imes 10^{-5}$      |
| 18 | $-5.61149517	imes10^{-4}$       | 48                    | $2.04694055 	imes 10^{-4}$      | 78            | $-3.94948537 	imes 10^{-5}$      |
| 19 | $9.06633351	imes10^{-4}$        | 49                    | $1.092~891~25 	imes 10^{-4}$    | 79            | $-3.59067012 	imes 10^{-5}$      |
| 20 | $1.96998543	imes 10^{-3}$       | 50                    | $1.85144590 	imes 10^{-5}$      | 80            | $-3.00174130 \times 10^{-5}$     |
| 21 | $2.62196298 	imes 10^{-3}$      | 51                    | $-6.11340629 	imes 10^{-5}$     | 81            | $-2.25881443 	imes 10^{-5}$      |
| 22 | $2.89300170 	imes 10^{-3}$      | 52                    | $-1.24961453	imes10^{-4}$       | 82            | $-1.43981869 	imes 10^{-5}$      |
| 23 | $2.83948651	imes10^{-3}$        | 53                    | $-1.70176902 	imes 10^{-4}$     | 83            | $-6.18258640 	imes 10^{-6}$      |
| 24 | $2.53321973	imes10^{-3}$        | 54                    | $-1.95836070	imes10^{-4}$       | 84            | $1.41864677	imes10^{-6}$         |
| 25 | $2.05221967\times10^{-3}$       | 55                    | $-2.026\:502\:29\times10^{-4}$  | 85            | $7.897~365~74 	imes 10^{-6}$     |
| 26 | $1.47313226	imes 10^{-3}$       | 56                    | $-1.92695992\times10^{-4}$      | 86            | $1.28991617	imes 10^{-5}$        |
| 27 | $8.65397089\times10^{-4}$       | 57                    | $-1.690\ 619\ 44\times 10^{-4}$ | 87            | $1.62305601 	imes 10^{-5}$       |
| 28 | $2.871\ 842\ 60\times 10^{-4}$  | 58                    | $-1.354\ 679\ 88\times 10^{-4}$ | 88            | $1.785\ 289\ 84 	imes 10^{-5}$   |
| 29 | $-2.169\ 887\ 80\times 10^{-4}$ | 59                    | $-9.58901614\times10^{-5}$      | 89            | $1.78650877 	imes 10^{-5}$       |
|    |                                 |                       |                                 |               |                                  |

**Table 2.** Renormalized perturbation expansion for the Hamiltonian (12) corresponding to the variable  $(1 - \kappa)$ , for  $\omega = 1$ , without the factor  $[\lambda/(1 - \kappa)]^{1/2}$ .

It is sometimes possible to extend the perturbative prediciton beyond these limits. One of the most common techniques for such an improvement in the convergence of a perturbation series is its renormalization. It consists of a nonlinear mapping of the original perturbation parameter ( $\lambda$  in the equations above) onto a more convenient one. There are many equivalent mappings (see, e.g., [3]). Here we consider one which has lately received detailed attention [18, 19] in connection to strong-coupling expansions. It changes the perturbation parameter  $\lambda \in [0, \infty]$  into  $\kappa \in [0, 1]$  according to

$$\lambda^{5/4} = \frac{\omega (1-\kappa)^{5/4}}{\kappa} \tag{11}$$

where the exponent  $\frac{5}{4}$  is appropriate for the Hamiltonian in (8), and  $\omega$  is a free parameter. Application of the scale transformation  $x \to x\sqrt{1-\kappa}$  to the Hamiltonian (8) leads to a new Hamiltonian

$$H^{(r)} = \sqrt{\frac{\lambda}{1-\kappa}} \left[ p^2 + (1-\kappa) \left( x^2 - \frac{\mathrm{i}x^3}{\omega} \right) + \frac{\mathrm{i}x^3}{\omega} \right]. \tag{12}$$

The algorithm used to calculate the original perturbation expansion also applies to the renormalized series generated by the Hamiltonian (12). The resulting expansion coefficients

are shown in table 2. After searching for an optimal value of  $\omega$  we chose  $\omega = 1$  in all our calculations.

**Table 3.** Ground-state energy calculated by the renormalized series and by direct numerical integration of the Schrödinger equation for a wide range of values of the coupling constant  $\lambda$ .

| λ   | к        | E(perturbative) | E(numerical)  |
|-----|----------|-----------------|---------------|
| 0.1 | 0.907 47 | 1.158 161 23    | 1.158 161 23  |
| 1   | 0.461 40 | 1.291 754 16    | 1.291 754 16  |
| 10  | 0.052 56 | 3.169 095 62    | 3.169 096 16  |
| 100 | 0.003 15 | 9.999 968 45    | 10.000 068 74 |

In the language of perturbation theory the  $\mathcal{PT}$  symmetry of our models and the related phenomenon of existence of real energies is reflected by the reality of the perturbative coefficients. The renormalized coefficients in table 2 exhibit a surprising oscillatory behaviour with an amplitude that decreases slowly with the perturbation order. We find this resummation-friendly behaviour remarkable, as it does not occur in similar expansions for real Hamiltonians [19].

By absolute value, the coefficients of the renormalized series decrease with the order more slowly than the coefficients of the standard expansion. Nonetheless, the former series exhibits better convergence properties because the new perturbation parameter  $(1 - \kappa)$  is limited to the interval [0, 1]. For an immediate check one may recall a direct numerical integration of Schrödinger equation for comparison. A small sample of such a test is given in table 3. For the ground-state energy and using the coefficients in table 2 we can see that the relative difference between the exact result and its remormalized perturbative approximant is always smaller than  $10^{-5}$ , even for huge values of the original unrenormalized coupling  $\lambda$ .

Clearly, renormalization is successful in this case. Let us emphasize that such an observation is non-trivial. Indeed, in contrast to the current experience with unitary equivalence of Hermitian operators, the explicit form of relationship between our present models (8) and (12) may only be characterized in terms of their  $\mathcal{PT}$ -symmetry in general.

#### Acknowledgments

RG and JR acknowledge financial support from the Dirección General de Investigación Científica y Tecnológica (DGICyT, Spain) under grant No PB97/1139. MZ acknowledges financial support from the Grantová agentura AV ČR, Praha under grant A 104 8602.

#### References

- [1] Itzykson C and Zuber J B 1980 Quantum Field Theory (New York: McGraw-Hill) p 464
- [2] Turbiner A and Ushveridze A G 1988 J. Math. Phys. 29 2053
   Le Gillou J C and Zinn-Justin J (ed) 1990 Large Order Behaviour of Perturbation Theory (Amsterdam: North-Holland)
- [3] Arteca A G, Fernández F M and Castro E A 1990 Large Order Perturbation Theory and Summation Methods in Quantum Mechanics (Berlin: Springer)
- [4] Guardiola R, Solís M A and Ros J 1991 Nuovo Cimento B 107 713
  Fernández F M and Guardiola R 1993 J. Phys. A: Math. Gen. 26 7169
  Janke W and Kleinert H 1995 Phys. Rev. Lett. 75 2787
  Weniger E J 1996 Phys. Rev. Lett. 77 2859
- [5] Bender C M and Wu T T S 1969 Phys. Rev. 184 1231

Reed M and Simon B 1978 Methods of Modern Mathematical Physics IV (New York: Academic)

- [6] Caliceti E, Graffi S and Maioli M 1980 Commun. Math. Phys. 75 51
   Bender C M and Milton K A 1997 Phys. Rev. D 55 R3255
   Bender C M and Milton K A 1998 Phys. Rev. D 57 3595
- [7] Bender C M and Boettcher S 1998 J. Phys. A: Math. Gen. **31** L273–7
- [8] Ushveridze A G 1994 Quasi-Exactly Solvable Models in Quantum Mechanics (Bristol: Institute of Physics)
- Bender C M and Boettcher S 1998 Phys. Rev. Lett. 80 5243
   Cannata F, Junker G and Trost J 1998 Phys. Lett. A 246 219
- [10] Bender C M and Turbiner A 1993 Phys. Lett. A 173 442
- [11] Alvarez G 1995 J. Phys. A: Math. Gen. 27 4589
- [12] Bessis D 1992 private communication to MZ
- [13] Simon B and Herbst I 1978 Phys. Rev. Lett. 41 67
- Bishop R F, Flynn M F and Znojil M 1989 Phys. Rev. A 39 5336
   Znojil M 1990 Czech. J. Phys. 40 1065
   Znojil M 1997 J. Phys. A: Math. Gen. 30 8771
- [15] Guardiola R and Ros J 1996 J. Phys. A: Math. Gen. 29 6461
   Fernández F M and Guardiola R 1997 J. Phys. A: Math. Gen. 30 7187
- [16] Znojil M 1996 Phys. Lett. A 223 411
- [17] Hille E 1969 Lectures on Ordinary Differential Equations (Reading, MA: Addison-Wesley) ch 11
- [18] Weniger E J 1996 Ann. Phys., NY 246 133
- [19] Skála L, Čížek J, Kapsa V and Weniger E J 1997 Phys. Rev. A 56 4471